



MAKING A DIFFERENCE

The Andromeda Zn-Cu prospect in the Albany-Fraser Orogenic belt: A HT SQUID discovery

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Andromeda Zn-Cu Discovery



A GSWA Exploration Incentive Scheme co-funded discovery

- Andromeda is a geophysical discovery and the first VMS prospect of significance in the Albany-Fraser Orogenic Belt.
- Identified by a MLEM survey over a structural ellipse.
- The conductive plate is 450m from surface.
- Andromeda has characteristics consistent with Besshi-style VMS deposits.
- Ore comprises pyrrhotite, pyrite, sphalerite and chalcopyrite.



IGO's Streamlined Asset Portfolio

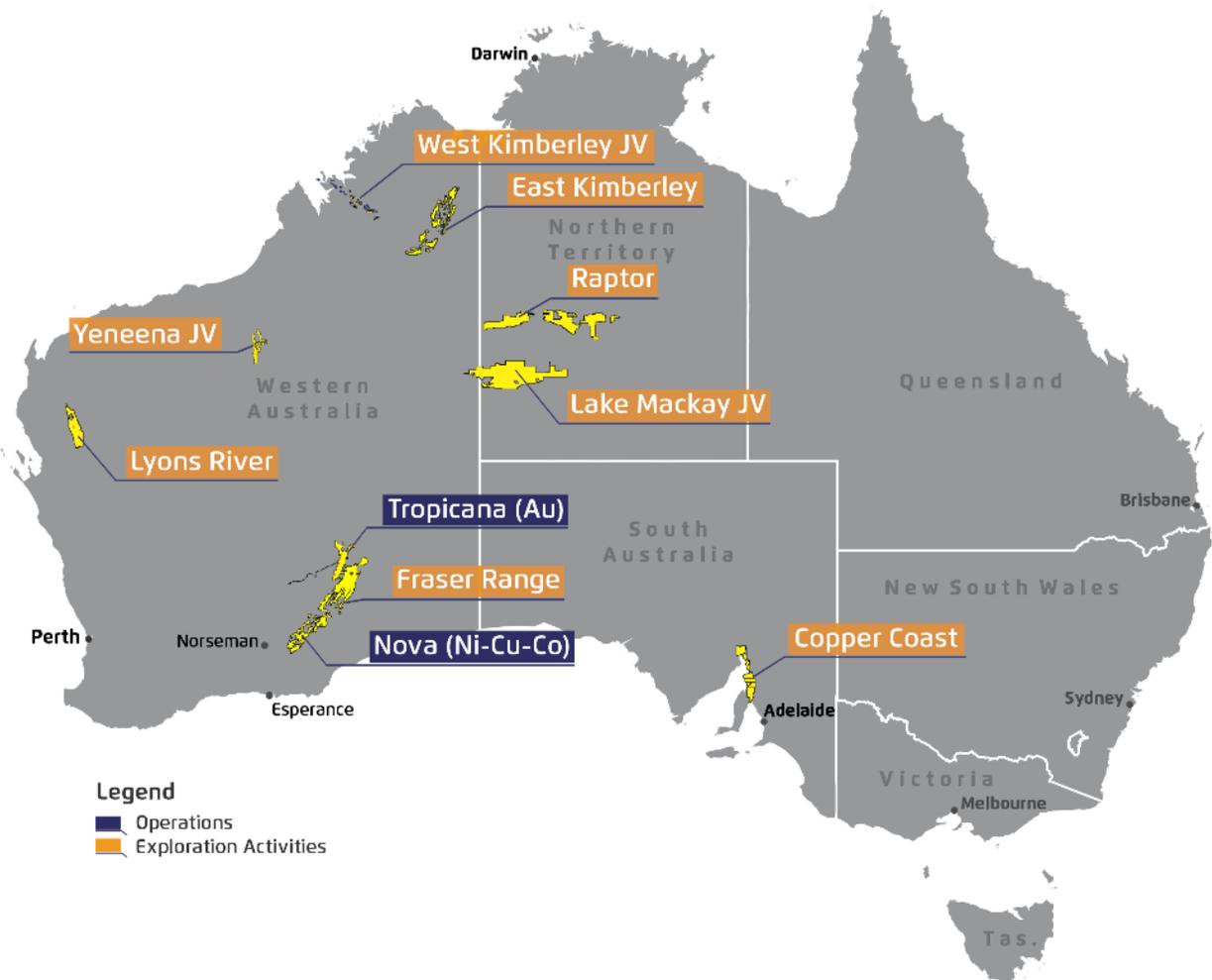


Belt-scale discovery opportunities critical for clean energy

Focus on orthomagmatic Ni-Cu and sediment-hosted Cu-Co mineralisation

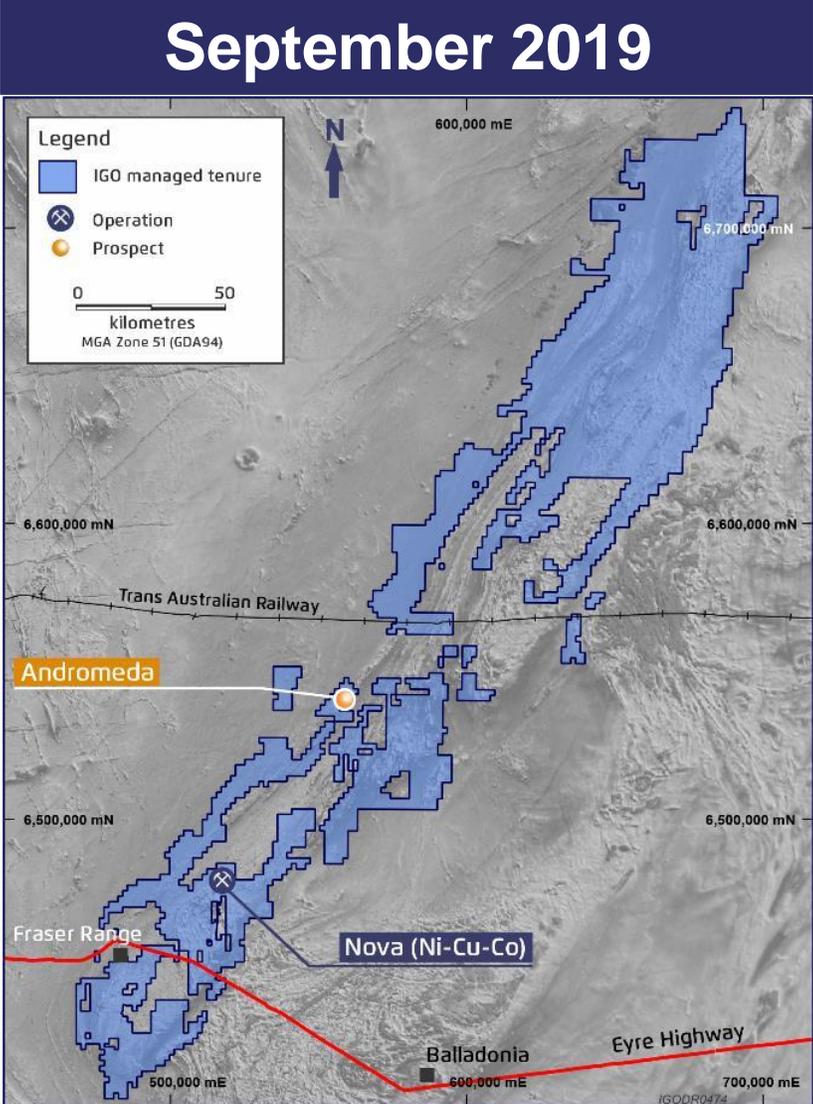
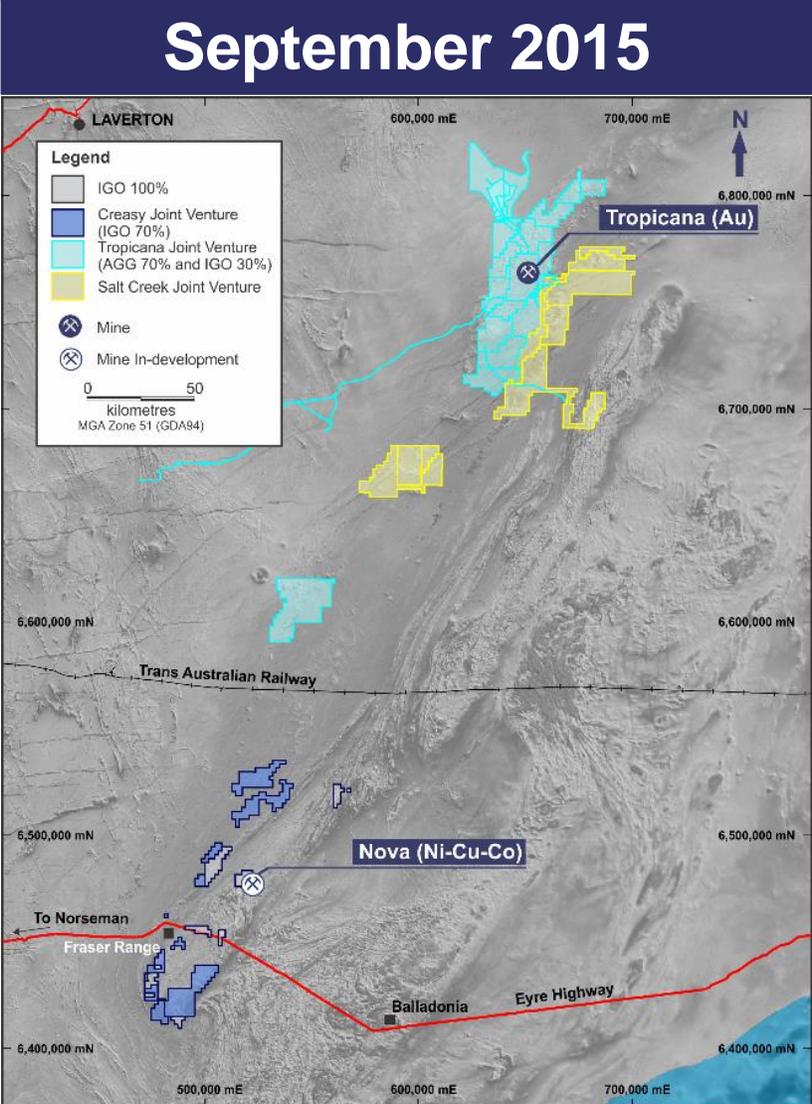
Belt scale land positions with the potential for significant base metals critical for clean energy

Large pipeline of new targets



Historic and current tenement position in AFO

Location of Andromeda Zn-Cu deposit in a Ni-Cu Belt

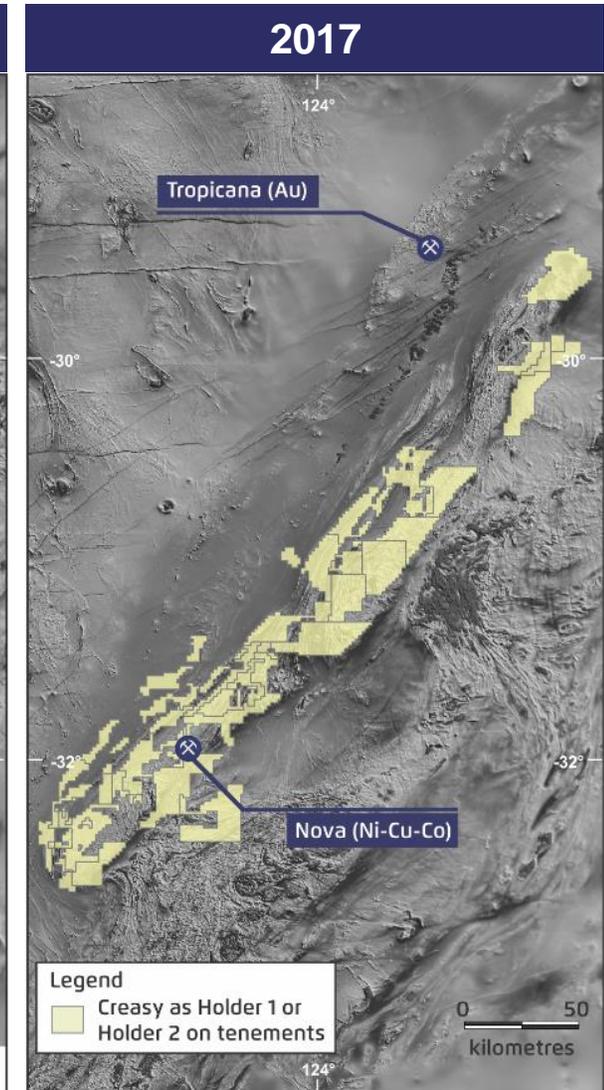
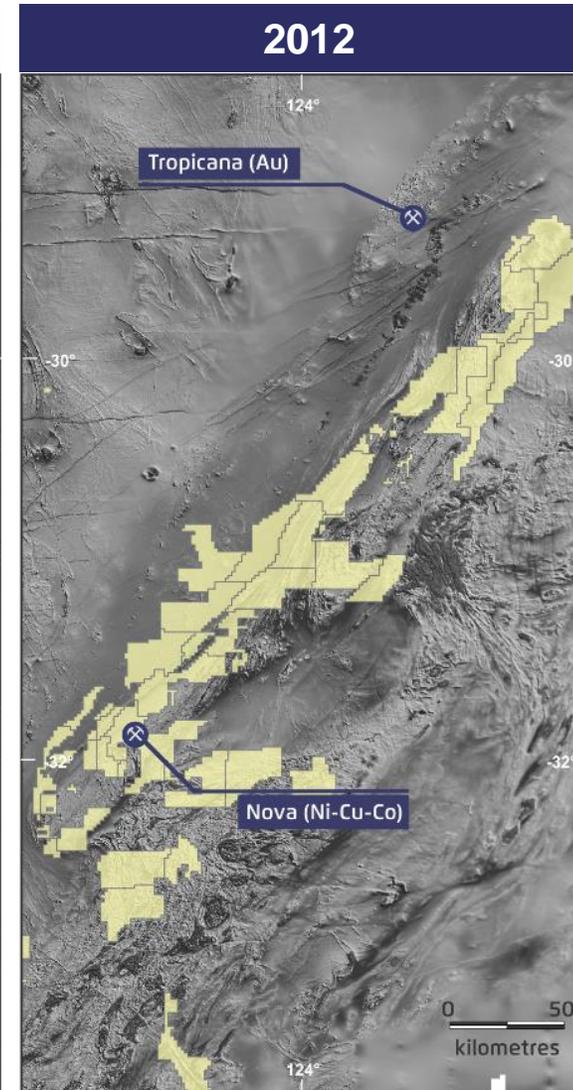
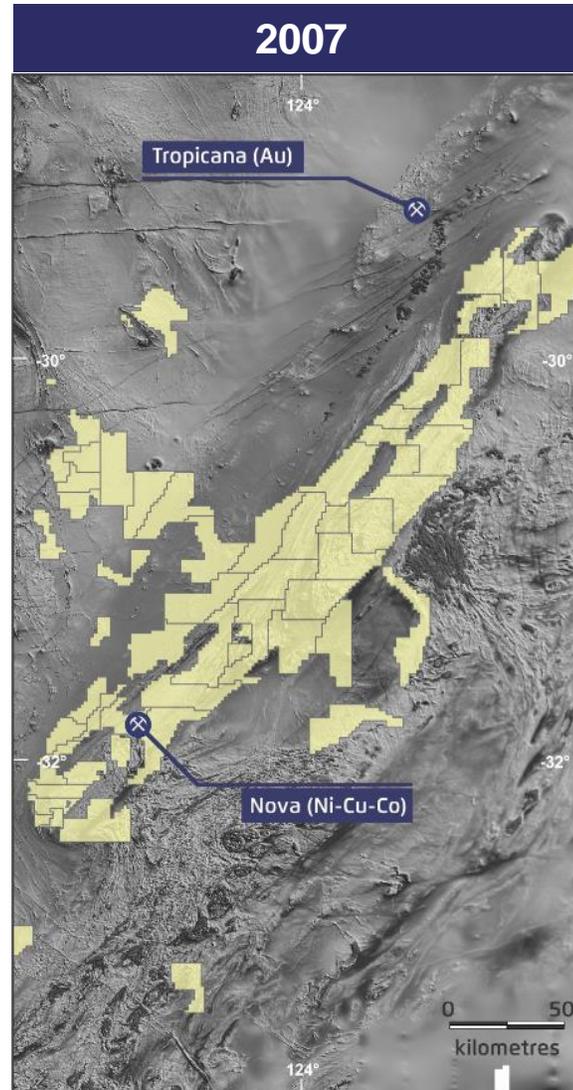


Albany Fraser Orogenic Belt



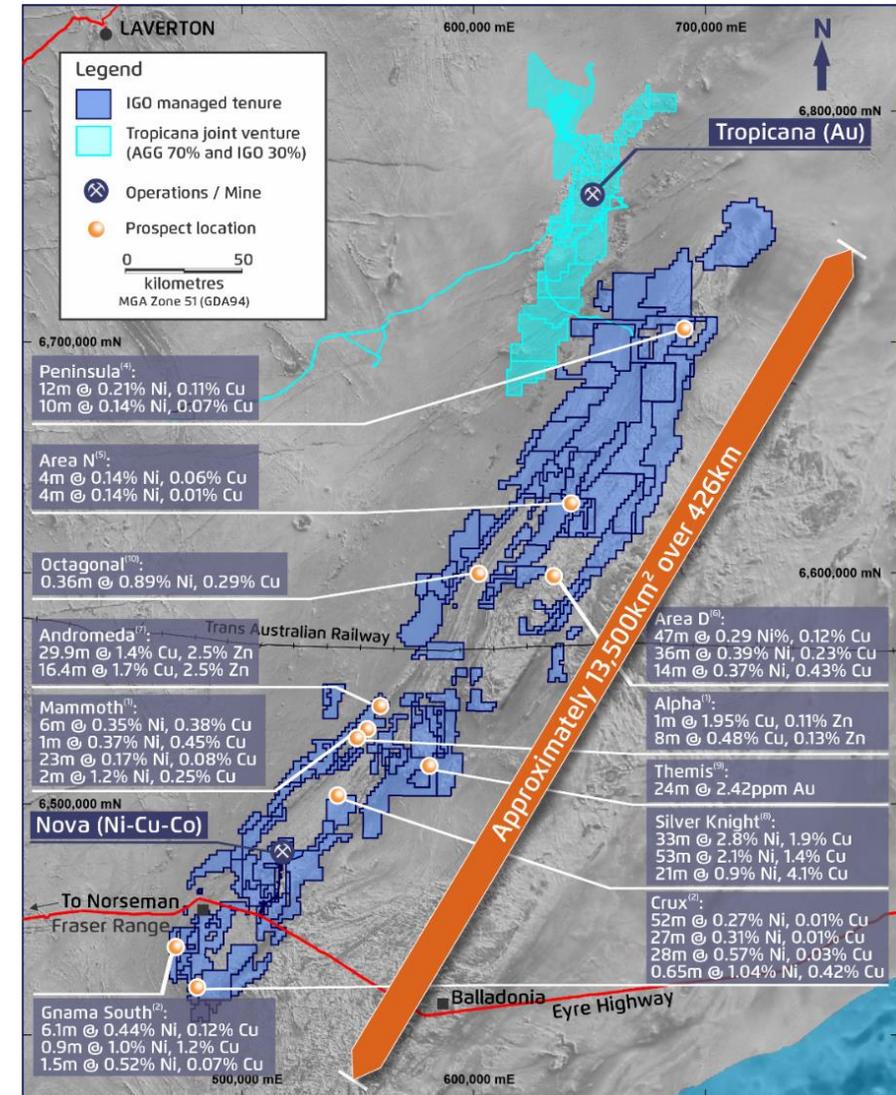
How explored is the Fraser Range?

- Creasy controlled companies have held and continue to hold large areas of tenement to today.
- A review of exploration activities in the Albany Fraser Belt reveals that exploration has been patchy and shallow.
- Creasy has done much work, but many opportunities remain.



Albany Fraser Orogenic Belt

A Ni-Cu tenement portfolio



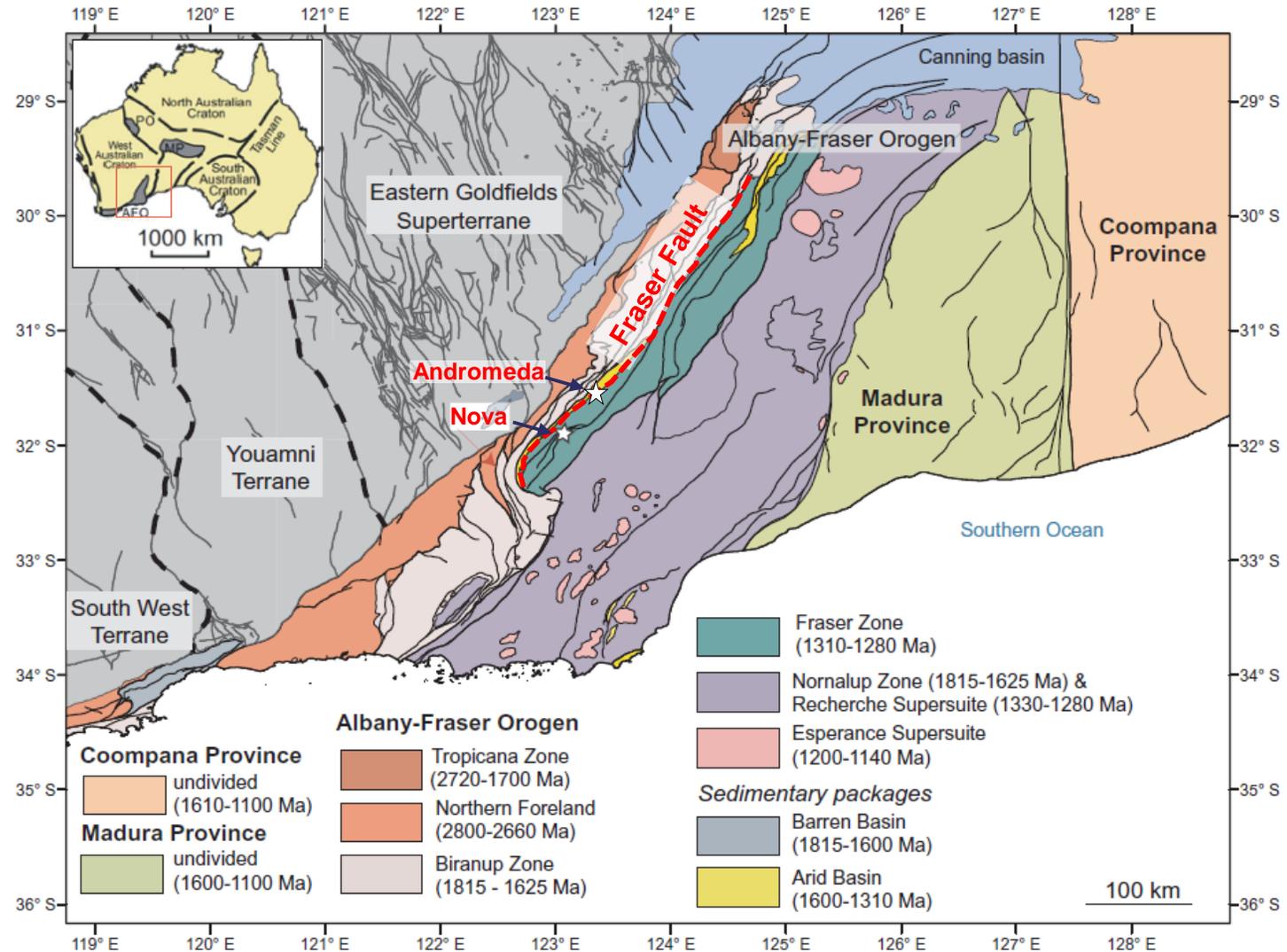
- Classic Minerals ASX Releases: 29 August 2013, 12 December 2013 and 17 December 2016
- Sirius Resources ASX Release June 2015 Quarterly and IGO ASX Releases: 26 July, 2018 and 20 February, 2019
- Enterprise Metals EIS Final Drilling Report to DMP: 25 July 2014
- Orion Gold ASX Release 17 March 2014
- Legend Mining ASX Release 6 June 2017, 12 January 2018, 12 April 2018, 9 July, 2019
- Creasy Group Application for M28/395 20 July, 2018
- Rumble Resource ASX Release, 1 July, 2019

Andromeda Zn-Cu Discovery

Regional Geology



- Andromeda is located within the Fraser Zone, a package dominated by granulite grade mafic rocks and lesser metasediments.
- The Fraser Zone is bound to the west by reworked Archean rocks, the Northern Foreland and rocks of oceanic affinity within the Madura Province to the east.
- Crustal architecture, plus geochemical data from within the Fraser Zone suggests a backarc or perhaps intracontinental rift setting (Glasson et al., 2019).
- Andromeda occurs along the margin between the 1600-1310 Ma Arid Basin and the 1310-1280 Ma Fraser Zone.

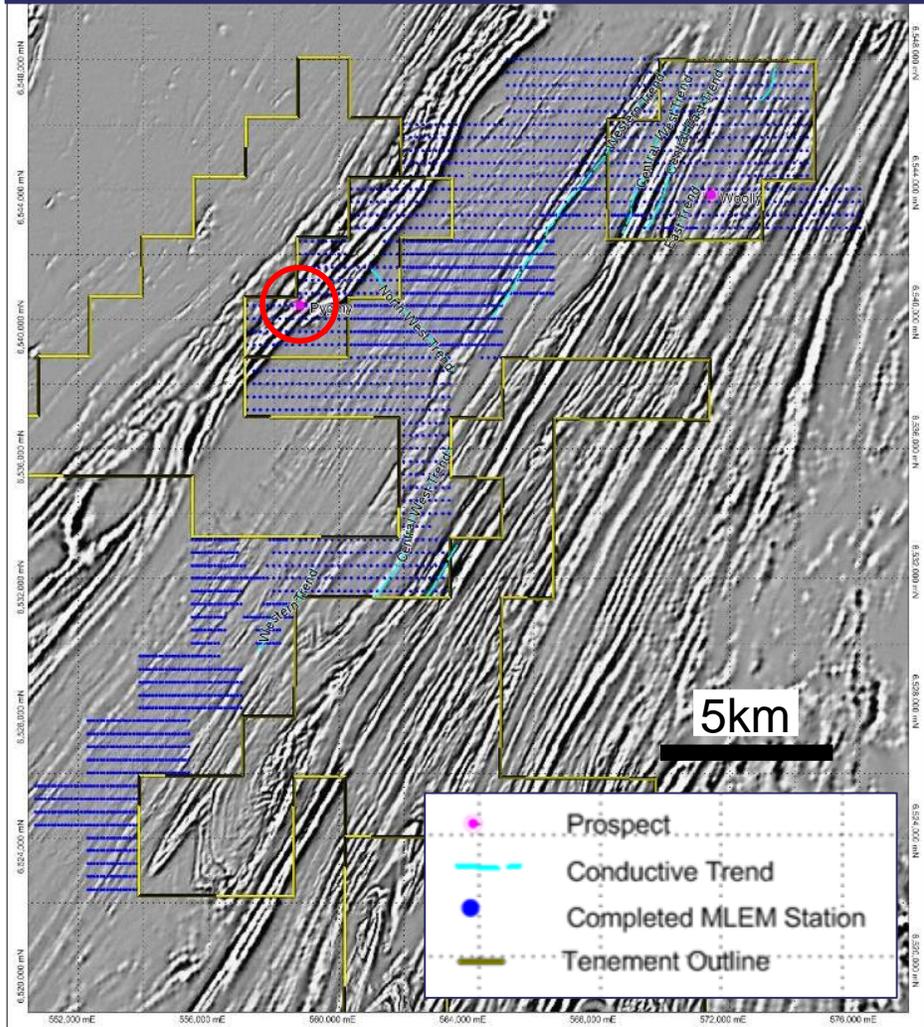


Andromeda Zn-Cu Discovery

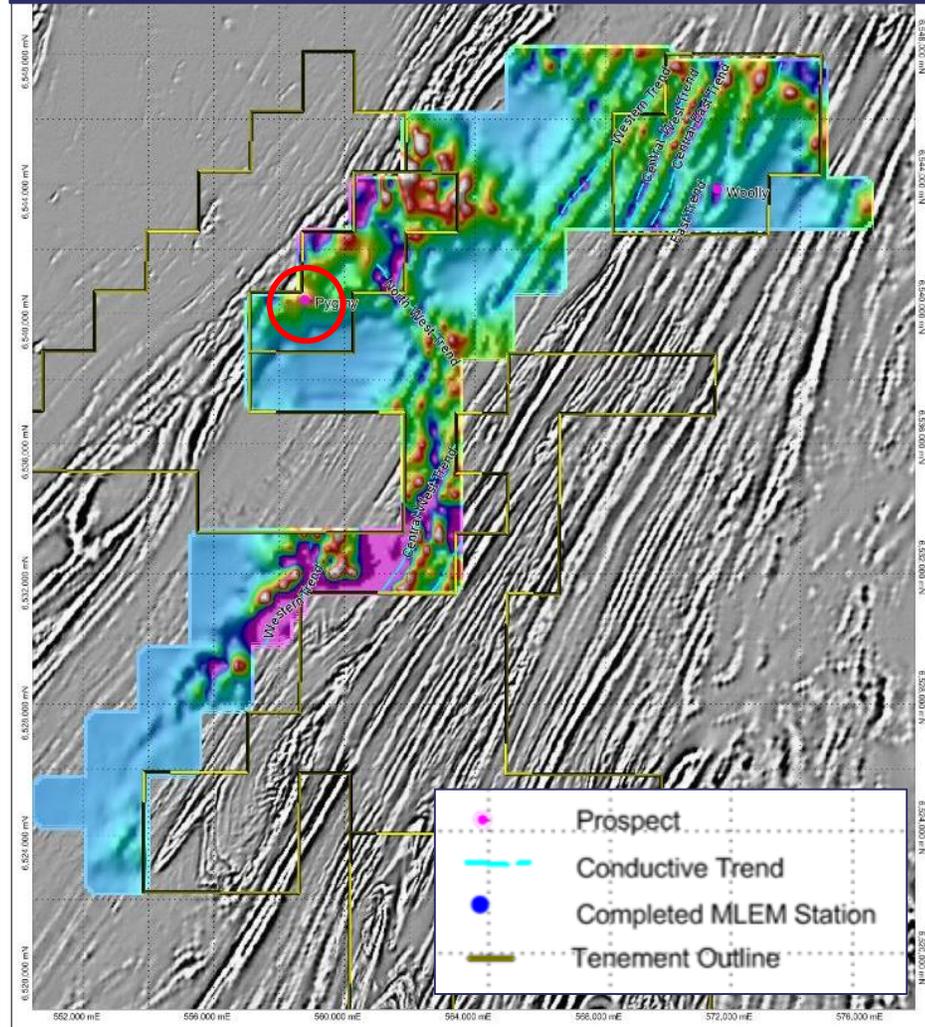
A deliberate combination of quality geophysics and geological curiosity



Tenements on TMI2VD, 400m station locations

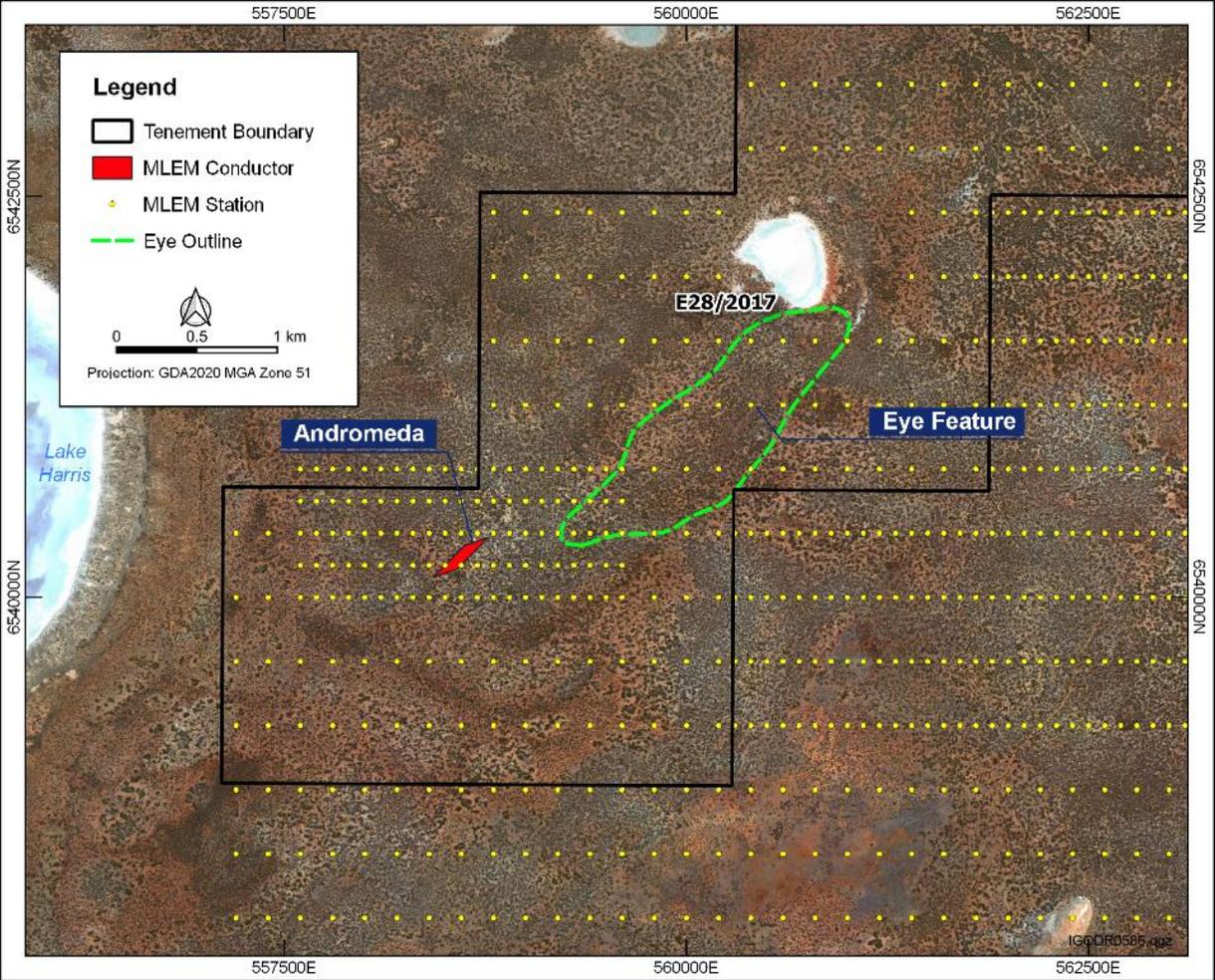
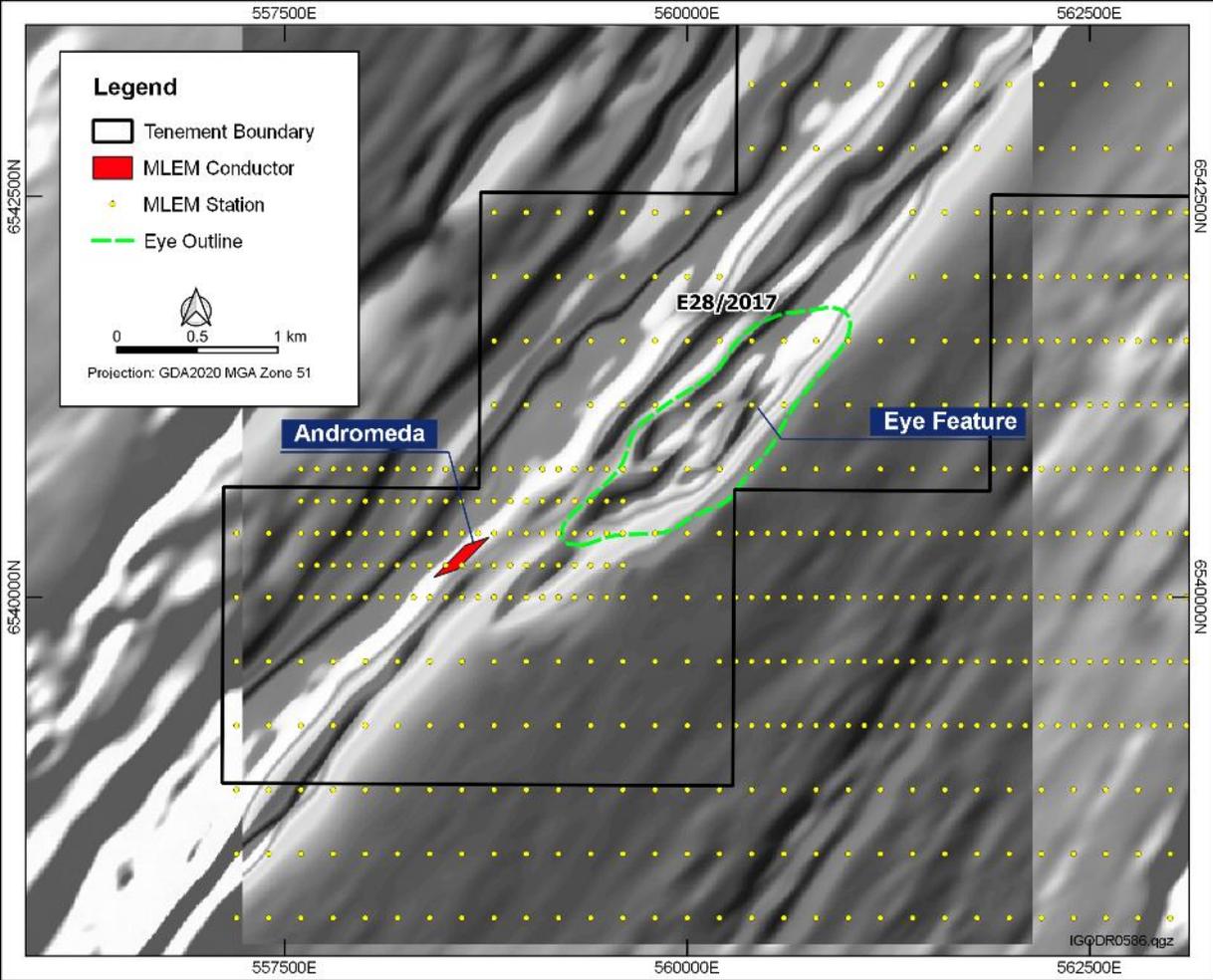


HT SQUID Z Ch25 showing conductive trends



Andromeda Zn-Cu Discovery

Geology masked by sand dunes, salt lakes and clay pans



Andromeda Zn-Cu Discovery

MLEM survey parameters



Configuration	Slingram to minimise IP and SPM effects
Loop Size	400m (200m infill)
Line Spacing	400m (200m infill)
Station Spacing	200m (100m infill)
Receiver system	Smartem24
Jessie Deeps HT SQUID	Bz (up), Bx (east), By (north)
Fluxgate	Bz(up), Bx (east), By (north)
Sensor Location	400m east of loop centre
Transmitter	Transmitter Technologies TTX
Effective current	~80A
Frequency	0.5Hz

Data repeatability for the HTS was $<0.02\text{pT/A}$ for the Z component.
Fluxgate repeatability was $\sim 0.1\text{pT/A}$ in the Z component.



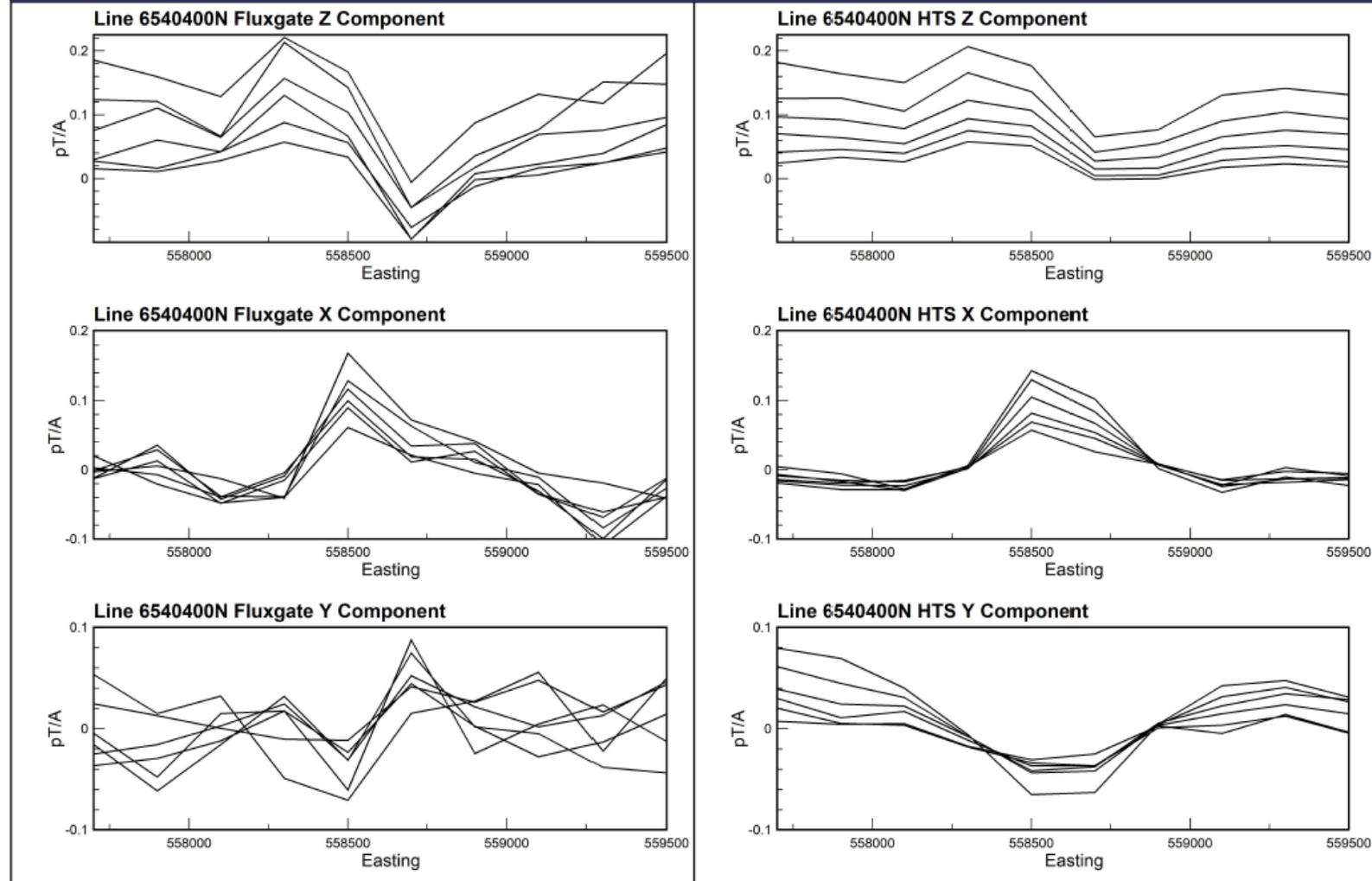
Andromeda Zn-Cu Discovery



Excellent detection at >450m

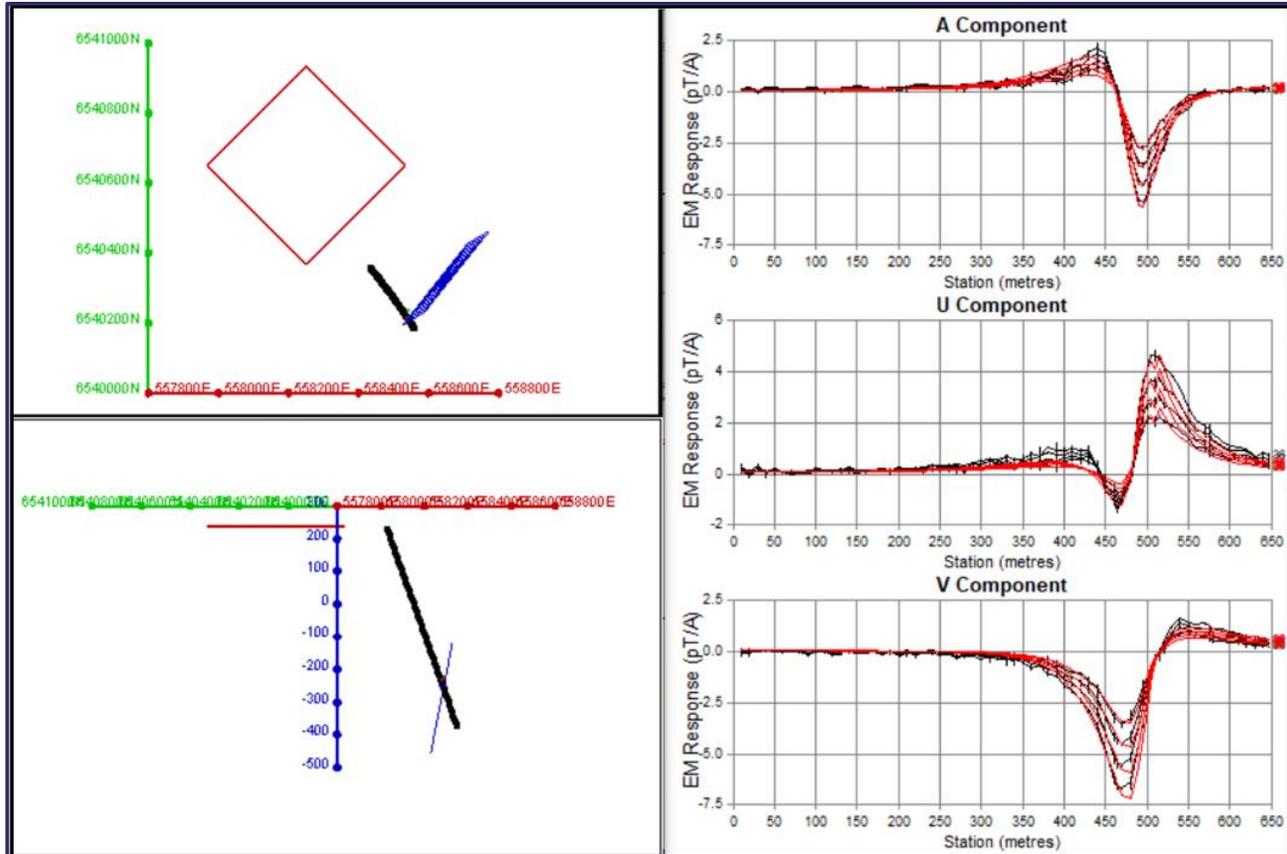
- The Andromeda conductor is seen as a late time anomaly with good responses in all three components of the HT SQUID readings.
- The response was characterised by an exponential decay with a time constant of approximately 80ms and interpreted to indicate massive sulphide mineralisation.
- Model parameters were best fit by a plate 400m long, 100m wide, 450m deep, dipping 80 degrees to the west and with a conductivity thickness of 4,200 Siemen.
- The fluxgate sensor also detected Andromeda, however the fluxgate signal was noisier.

MLEM profiles over Andromeda; Fluxgate (left) and HT SQUID (right)



Andromeda Zn-Cu Discovery

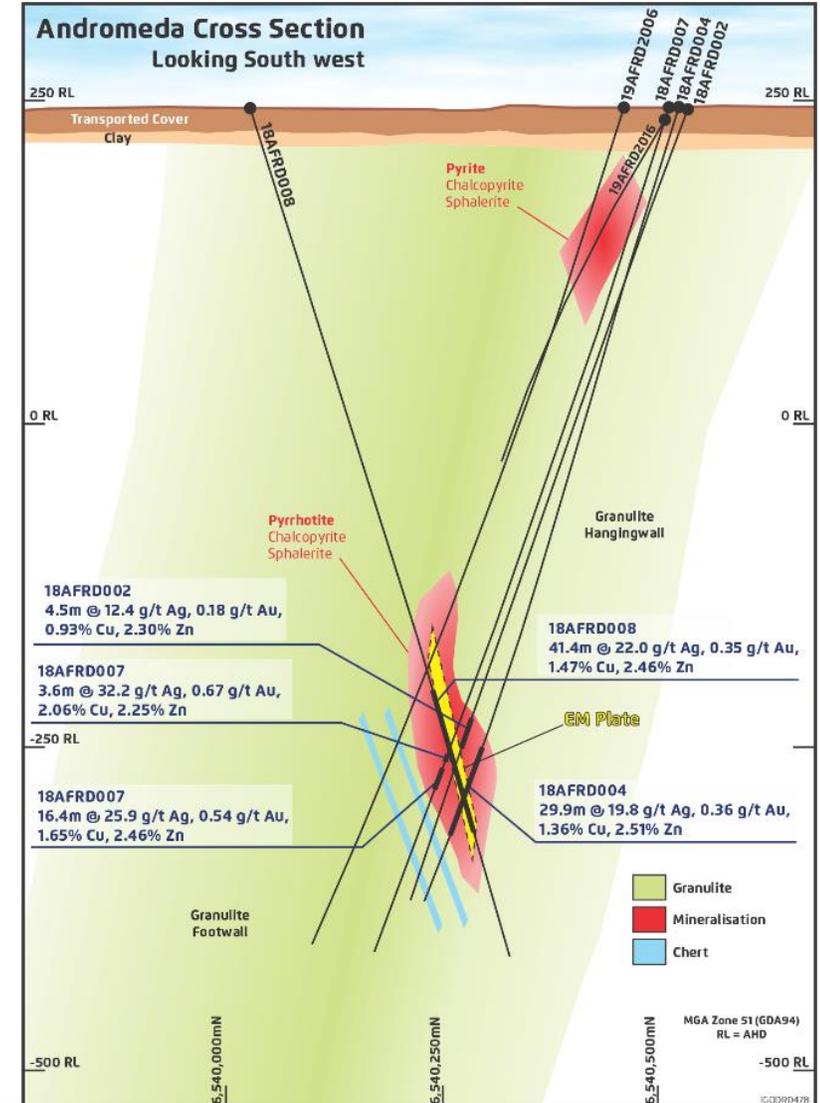
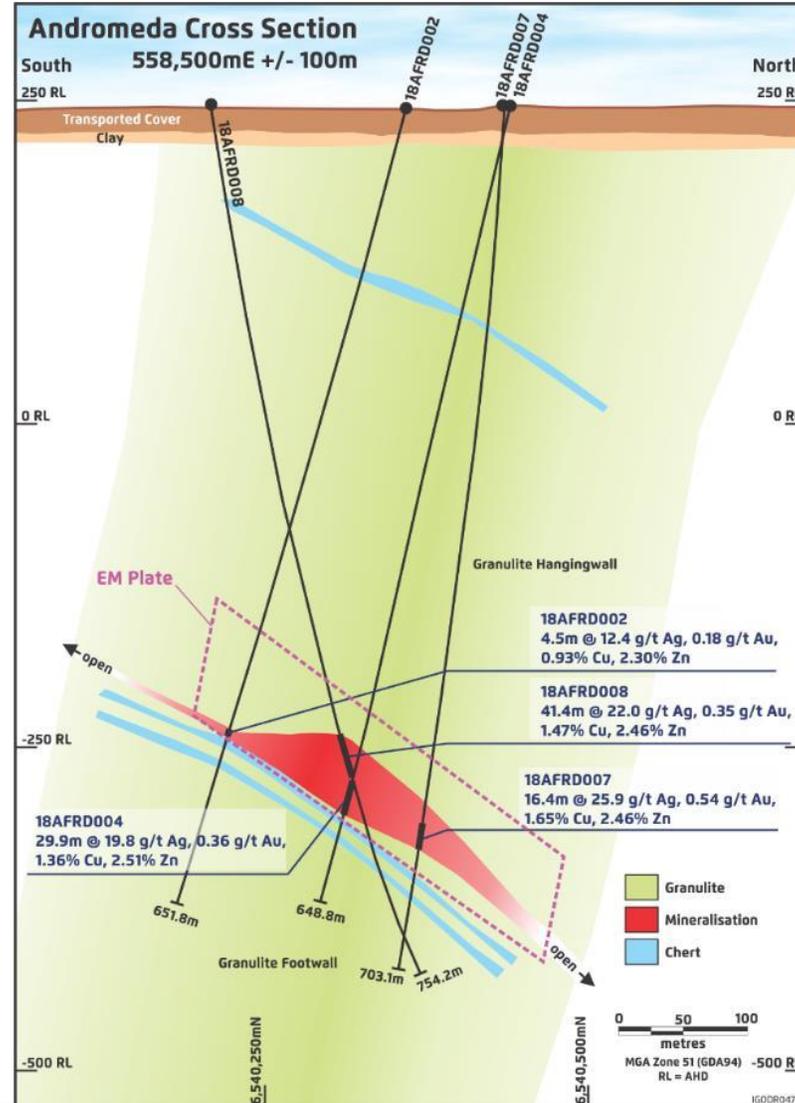
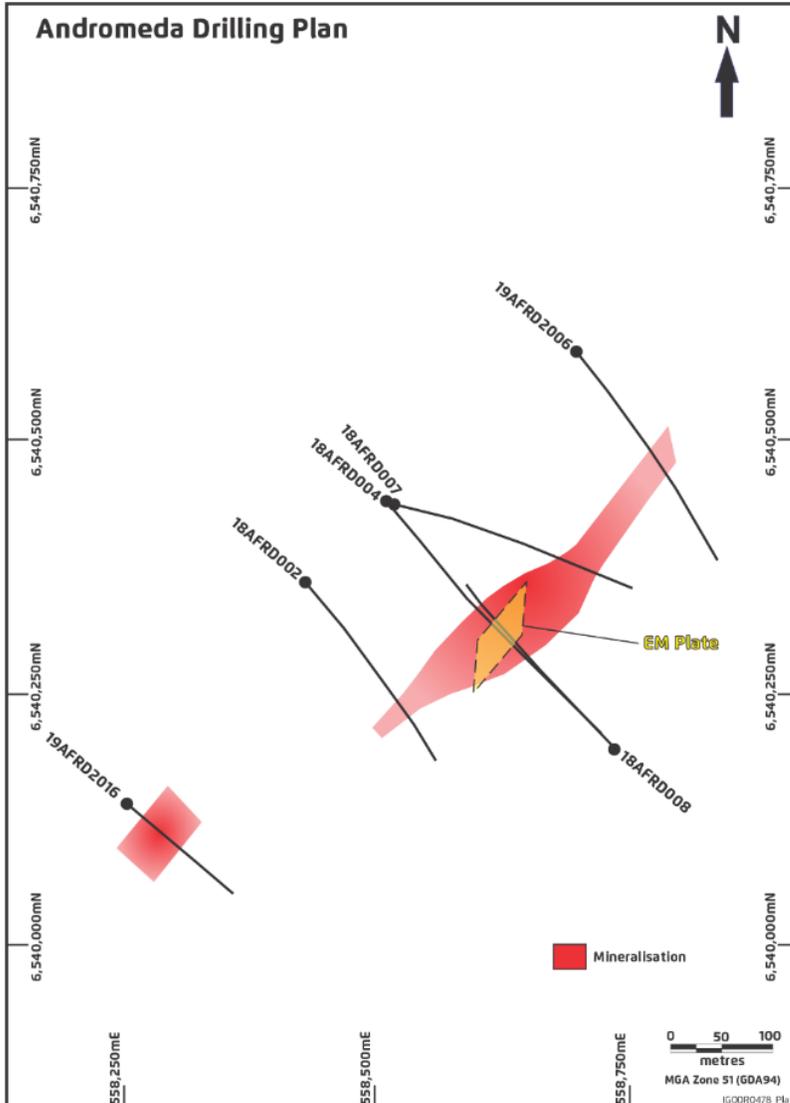
DHEM results



- Drill hole 18AFRD002 intersected the EM target ~80m from the planned pierce-point.
- 4.5m of pyrrhotite, pyrite, sphalerite and minor chalcopyrite was drilled from 510.5m downhole.
- 18AFRD002 was logged to 648m with a DigiAtlantis B-field EM probe, coupled with a transmitter with an output of 50A.
- Three conductors were identified with the strongest at ~515m, with time-constants in the order of ~200ms.
- A 400m x 100m, 24,000S off-hole conductor projected to align with the 5m intersection of pyrrhotite-dominant sulphides was modelled.

Andromeda Zn-Cu Discovery

Sections through the ore, including a new lens



Andromeda Zn-Cu Discovery



A high level look at the ore

- Ores show pervasive, evenly distributed, durchbewegung textures characterized by lithic fragments of (i) mafic host rocks and (ii) polycrystalline quartz/chert grains.
- No metal zonation across ore body; plastic deformation and 'movement' of ore within original host rock package.
- Very small stockwork zone underlies the deposit.

Top two photos:
Typical breccias of quartz/chert clasts in a pyrrhotite-chalcopyrite matrix.
Bottom two photos:
Cpy-Sph-rich matrix with >1cm dark silica/chert or mafic clasts.



Geochemical results from mineralized intersections through Andromeda ¹							
Hole ID	From (m)	To (m)	Width (m)	Cu %	Zn %	Au ppm	Ag ppm
18AFRD002 ¹	510.50	515.00	4.5	0.93	2.30	0.18	12.4
18AFRD004 ¹	548.10	578.00	29.90	1.36	2.51	0.36	19.8
18AFRD007 upper ¹	539.06	542.69	3.63	2.06	2.25	0.67	32.2
18AFRD007 main ¹	547.61	564.00	16.39	1.65	2.46	0.54	25.9
18AFRD008 #1 ¹	435.02	438.54	3.52	2.51	2.32	0.38	31.6
18AFRD008 #2 ¹	499.57	504.24	4.67	0.63	1.92	0.38	12.1
18AFRD008 #3 ¹	531.18	572.54	41.36	1.47	2.46	0.35	22.0
18AFRD008 #4 ¹	576.70	584.55	7.85	0.77	2.11	0.68	13.3

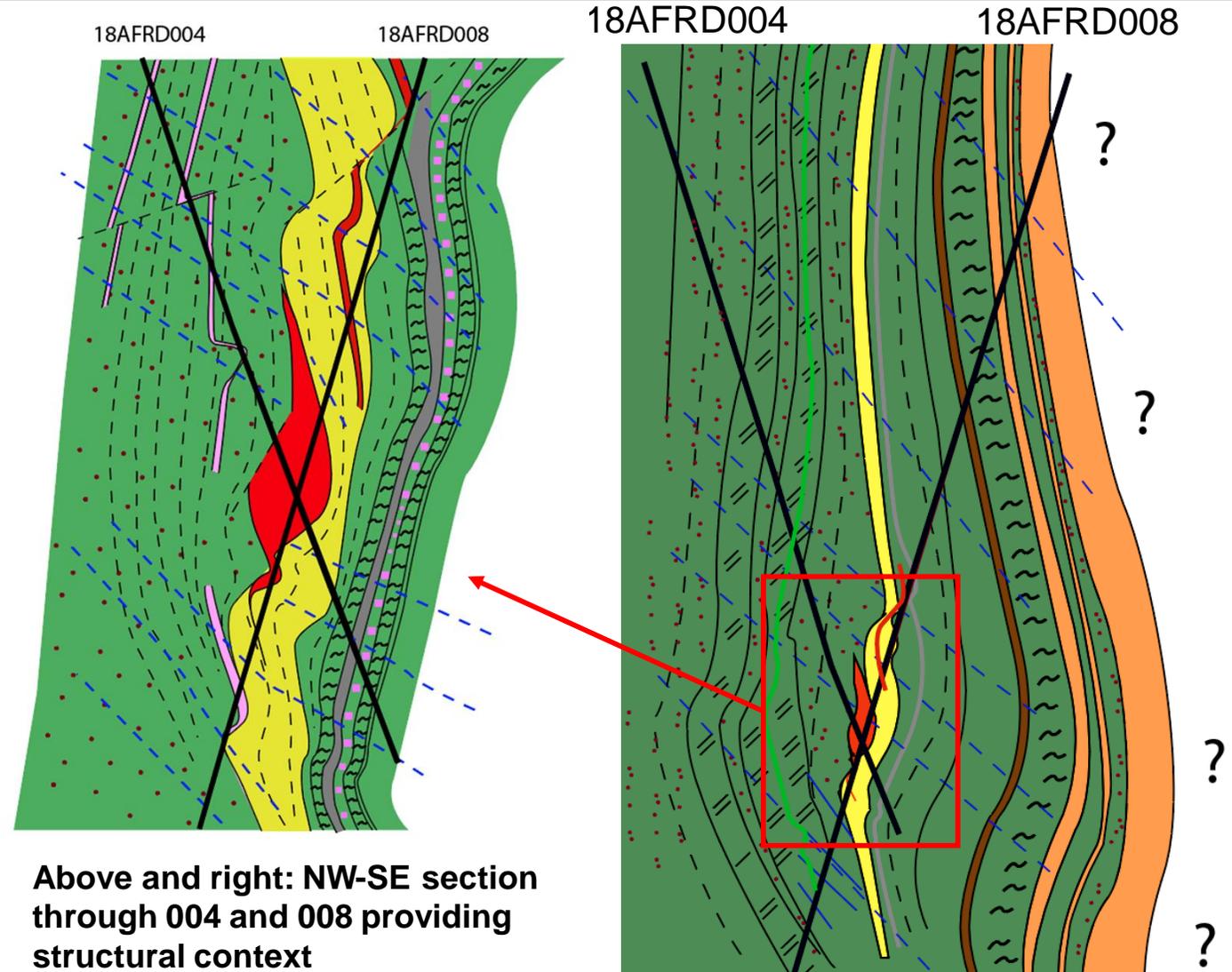
¹) Refer to IGO ASX releases dated 26 July, 2018 and 20 February 2019: Annual Update of Exploration Results, Mineral Resources and Ore Reserves

Andromeda Zn-Cu Discovery

Host rock geology to mineralisation

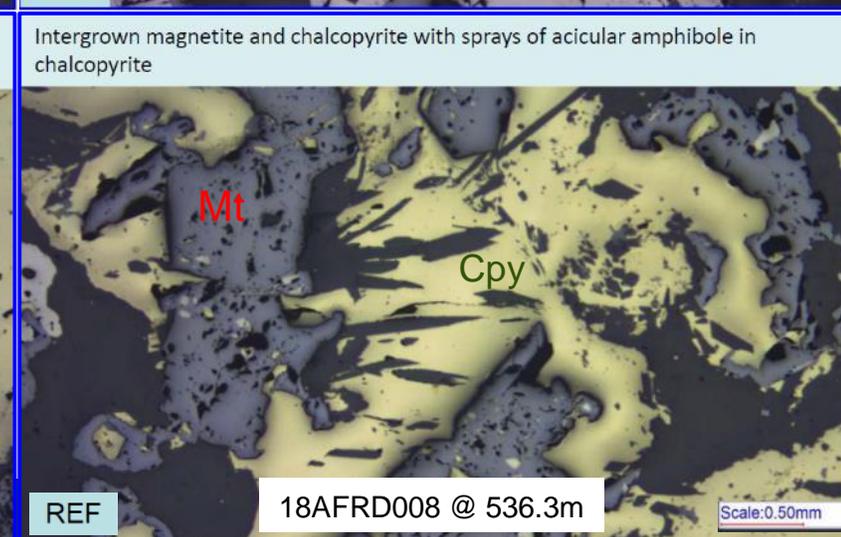
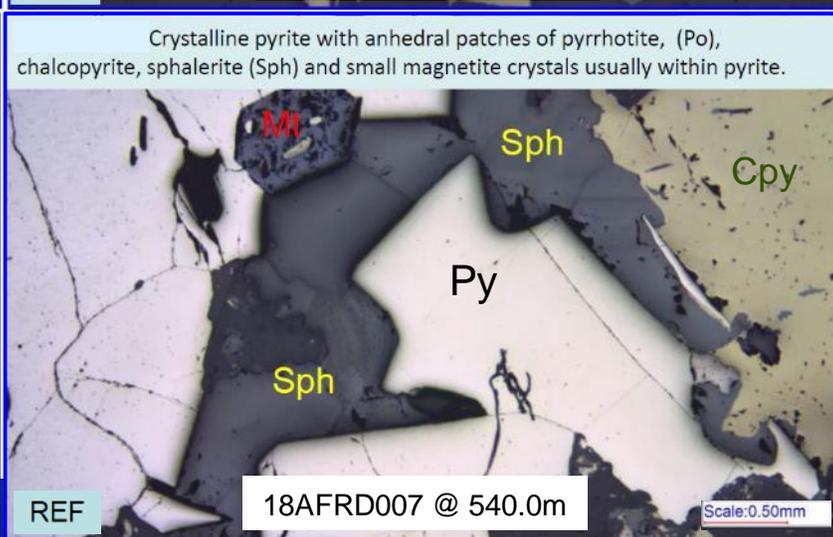
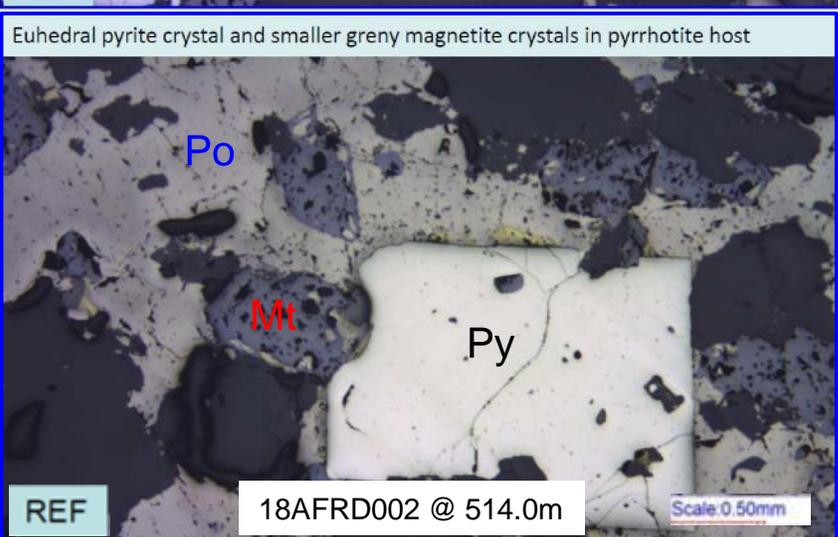
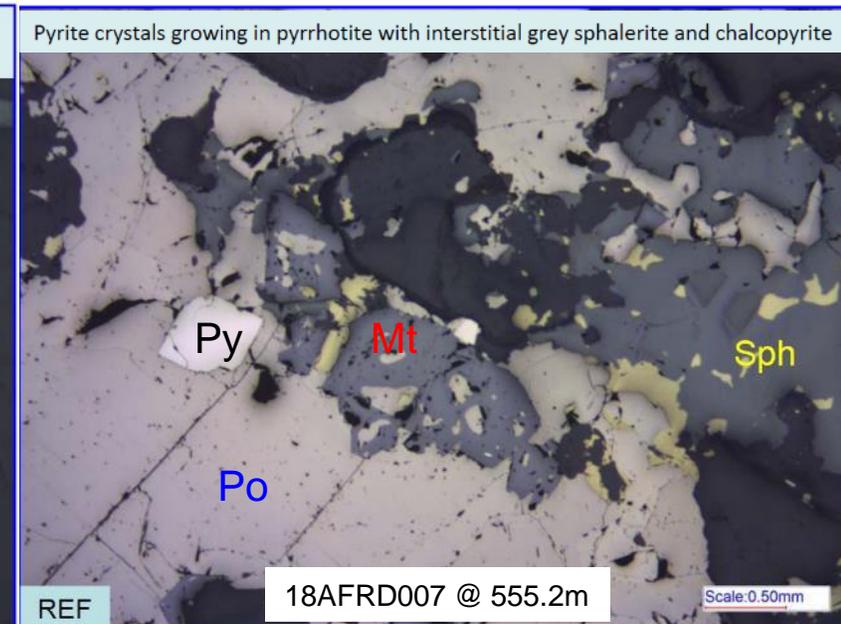
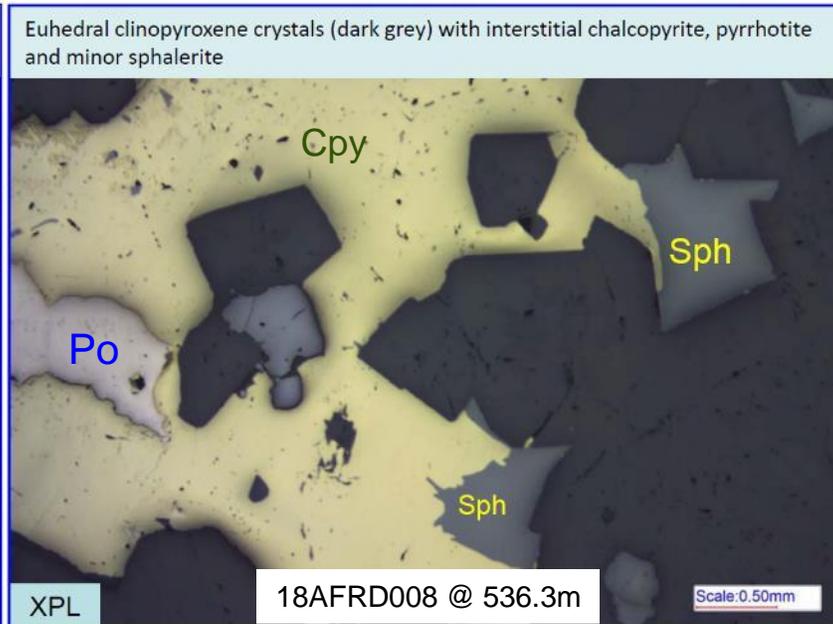
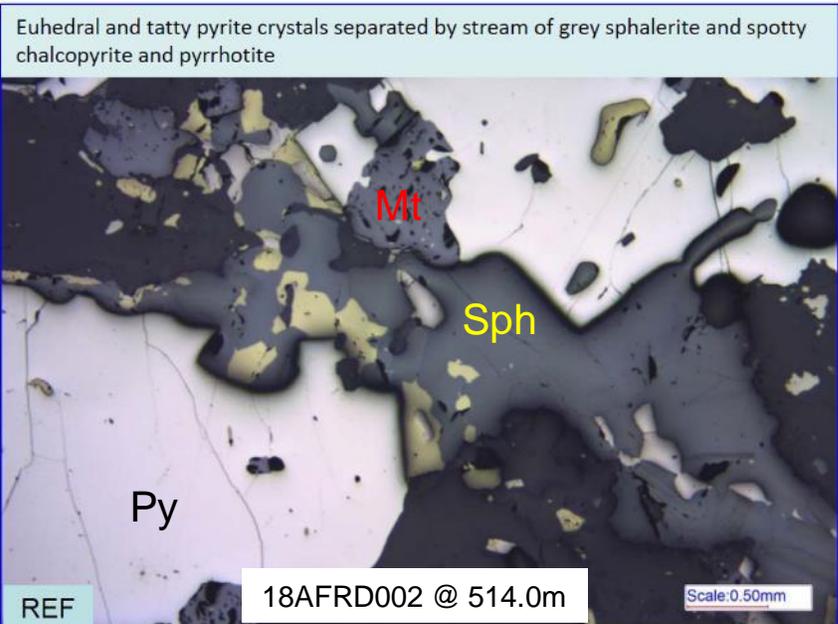


- Steeply dipping to subvertical mafic-dominated stratigraphy comprising garnetiferous mafic granulites (± 2 pyroxene), some hosting mottled to massive amphiboles.
- Sediments and orthogneisses only observed in the upper part of 18AFRD008, comprising strongly foliated, graphite-bearing metasediments of pelitic origin folded about asymmetric folds.
- Mineralisation displays sheared contacts – offsets undetermined
- Structural interpretation suggests that EM modelling is averaging two mineralised lenses – untested residual targets exist.



Andromeda Zn-Cu Discovery

Petrographic analysis of the ore



Andromeda Zn-Cu Discovery



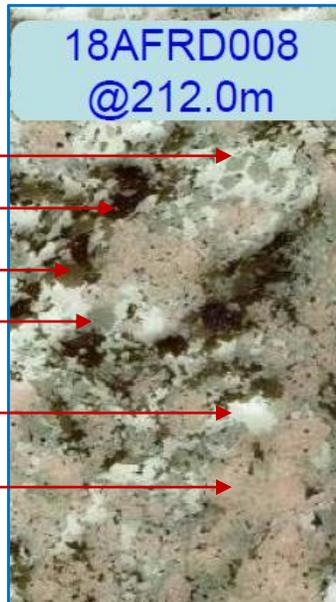
Mineralisation in a metabasalt – sill sequence

- Dominated by mafic granulites; divided into Group 1 and Group 2 on the basis of Si, Ti and Nb concentrations.
- Group 1: Mafic tholeiitic volcanic rocks, with minor interbedded metasedimentary rocks and leucosomes.
 - Zr/Ti and Nb/Y are consistent with basalts erupted during advanced stages of rifting of continental crust prior to breakup.
- Group 2: Basaltic magma that is considered to represent sills emplaced into the Andromeda rock package.

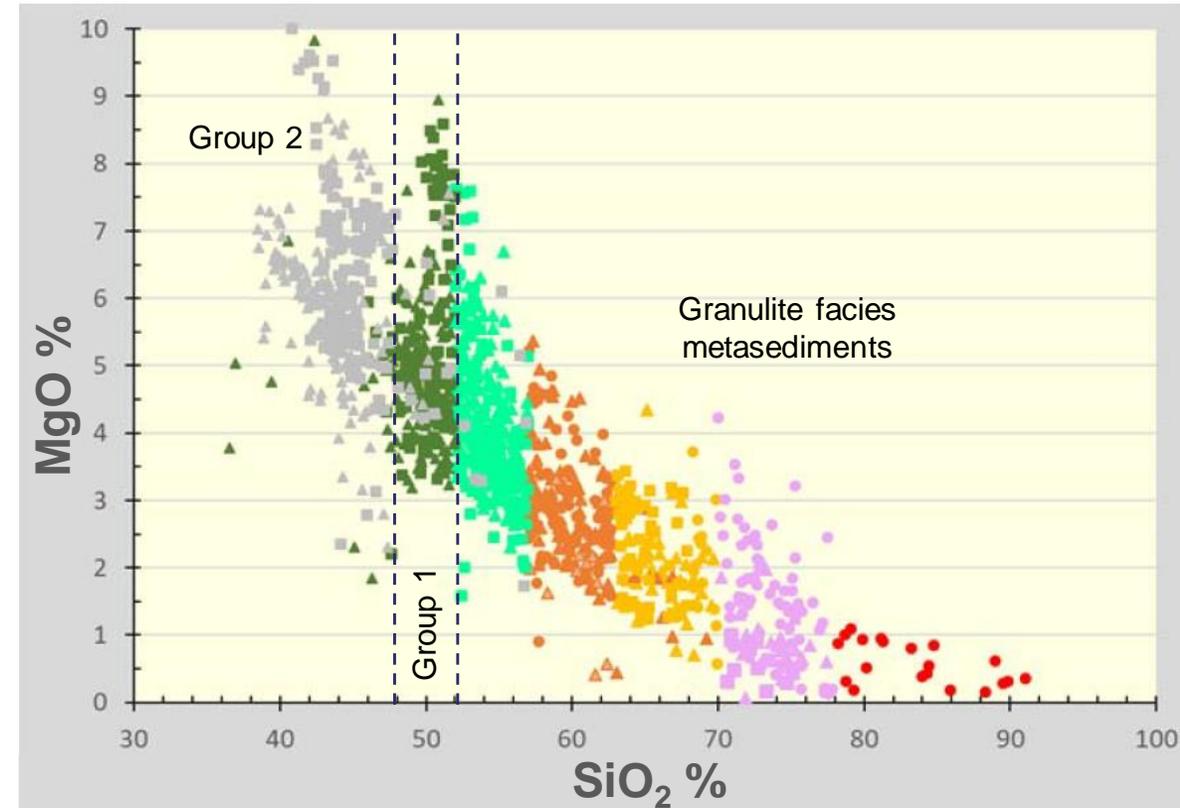
Group 1 Mafic Granulites: garnet-bearing gabbros, to garnet-bearing norite.
~46-52% SiO₂, Ti-, Nb-poor



Hypersthene
FeTi oxides
Ti hornblende
Augite
Plagioclase
Garnet



Group 2 Mafic Granulites: Ti-rich, garnet bearing gabbros, to garnet-bearing norite.
mostly <<48% SiO₂, Fe-, Ti-, Nb-rich

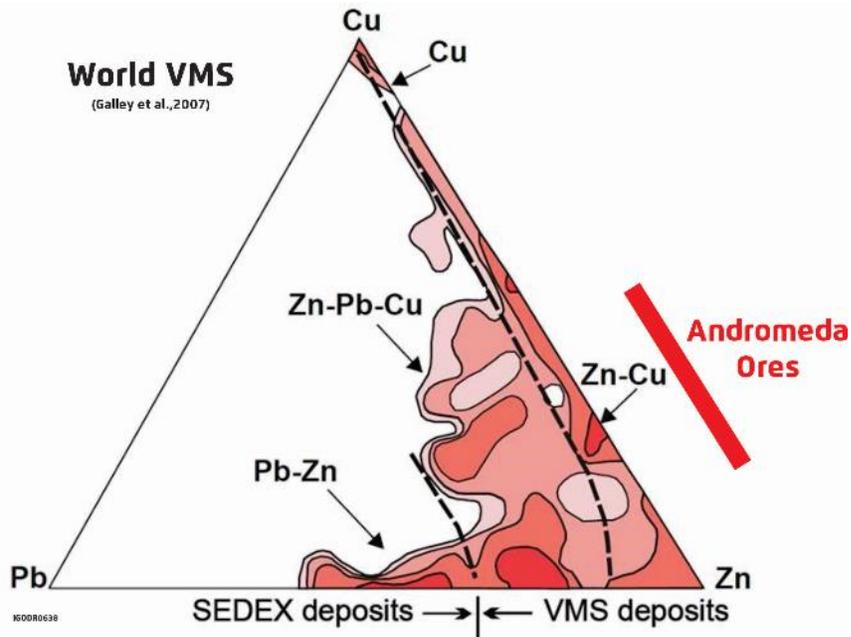


Andromeda Zn-Cu Discovery

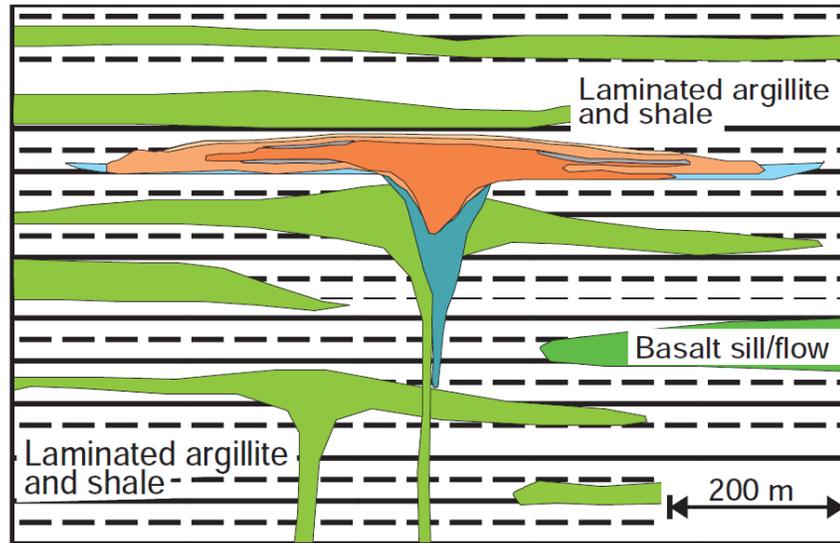


A siliciclastic-mafic, or Besshi-type VMS deposit

- The main ore zone comprises pyrrhotite, pyrite, sphalerite, magnetite, chalcopyrite and rare galena.
- 'Cu ratios' ($100\text{Cu}/(\text{Cu}+\text{Zn})$) are largely in the range 35-50, with no Pb.
- 'Zn ratios' ($100\text{Zn}/(\text{Zn}+\text{Pb})$) are >95 ; Andromeda is a Zn-Cu VMS deposit.
- Host-rock package is dominated by mafic lavas and sills interlayered with deep water sedimentary rocks.
- There is no preserved metal zonation across the orebody and a footwall stringer zone appears to be absent. Besshi-style deposits in Japan lack stockworks beneath the ore lenses and hydrothermal alteration zones are narrow.



Range of base metal compositions of VMS deposits from Galley et al. (2007) showing the range of $\text{Cu}/(\text{Cu}+\text{Zn})$ values for Andromeda ores with $>20\%$ S.



- | | |
|--|--------------------------|
| Pyrrhotite-pyrite-magnetite transition zone | Chert-carbonate-sulphide |
| Pyrrhotite-pyrite-chalcopyrite zone | Pyrite-sphalerite zone |
| Pyrrhotite-chalcopyrite-pyrite-sphalerite stockwork zone | Massive pyrite zone |

PELITIC-MAFIC

Canadian grade and tonnage

Average 34.3 Mt
Median 148 Mt

1.6% Cu
2.6% Zn
0.36% Pb
29 g/t Ag
<0.9 g/t Au

- The Andromeda VMS mineralisation is characterised by a pyrrhotite > pyrite >> sphalerite, chalcopyrite and magnetite assemblage, making it easily detected by both fluxgate and HT SQUID EM sensors.
- Metamorphism obliterated primary ore textures, homogenised the Cu/Zn ratio and blended the ores with their cherty metasedimentary and mafic host rocks.
- The host rock succession is dominated by Group 1 oxide-poor and Group 2 oxide-rich mafic granulites.
- Group 1 mafic granulites have rift tholeiite compositions consistent with basalts erupted during the advanced stages of rifting of continental crust just before breakup.
- Group 2 mafic granulites are interpreted as ferrogabbro sills that were intruded into the volcanic package.
- The Andromeda mineralisation type, ore grade and host rock package closely resembles that of Besshi-type deposits.
- The discovery of Andromeda opens a new search space in the AFO for Zn-Cu mineralisation along with the orthomagmatic Ni-Cu-Co mineralisation already known to exist.



Acknowledgements

- Geological Survey of Western Australia are thanked for co-funding drill hole 18AFRD002 through the EIS grant system.
- Jacob Paggi proposed extending the original MLEM survey west to include the Andromeda “eye” feature that lead to the identification of the EM conductor.
- The IGO Exploration Team.





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